

PAN-EUROPEAN CONSORTIUM FOR AVIATION SPACE WEATHE USER SERVICES





MUF(3000) nowcasting as operation space weather product



ISTITUTO NAZIONALE DI GEOFISICA E VULCANOLOGIA D. Sabbagh, C. Scotto, P. Bagiacchi Istituto Nazionale di Geofisica e Vulcanologia, via di Vigna Murata 605, 00143 Rome, Italy

e-mail address: dario.sabbagh@ingv.it

Pan-European Consortium for Aviation Space weather User Services (PECASUS) is one of the three global Space Weather Centers for aviation space weather user services designed by the International Civil Aviation Organization (ICAO).

The MUF(3000) nowcasting is one of the operational space weather products inserted in PECASUS. A mapping procedure is then applied to the European stations providing MUF(3000) nowcasting over the whole area. This procedure consists in upgrading the IRI-CCIR model using available real-time measurements and the Ordinary Kriging method for spatial interpolation.

Mapping method			Table 1
The value of the variable <i>z</i> for a given time is spatially interpolated using the Ordinary Kriging method, being:	Station	Longitude (°E)	Latitude (°N)
The value of the variable 2 for a given time is spatially interpolated using the Orumary Kriging method, being.	Chilton	-00.6	51.5
$MUF(3000)_{[meas]}(x_i) - MUF(3000)_{[med]}(x_i)$	Gibilmanna	14.0	37.9
$Z(\boldsymbol{x}_i) = \frac{1}{ \mathbf{M} \mathbf{E}(2000) - (\mathbf{x}_i) } $	Testinessee	10 /	

$MUF(3000)_{[me]}(x_i)$ /	Juliusruh	-13.4	54.6
where MUE(2000) is the value obtained by the IDI CCID model and rearranhic coordinates of the <i>i</i> -th	Moscow	37.3	55.5
		-12.5	41.8
$MUF(3000)_{[me]}(x_i)$ where MUF(3000)_{[med]} is the value obtained by the IRI-CCIR model, and x_i are the geographic coordinates of the <i>i</i> -th ionospheric station where measurements are available (see Table 1). The interpolation is performed in the region A of longitude 12°W-45°E and latitude 32°N-72°N, over a grid with spatial resolution 0.5°x0.5°. Then, upgraded MUF(3000) maps are produced over the interpolation grid as: $MUE(2000)(w) = (1 + z(w)) - MUE(2000) = (w) = x \in \operatorname{grid} in A$ $MUE(2000)(w) = (1 + z(w)) - MUE(2000) = (w) = x \in \operatorname{grid} in A$	52.2		
		-24.0	38.0
		14.6	50.0
	Dourbes	04.6	50.1
$MUF(3000)(x) = (1 + z(x)) \cdot MUF(3000)_{[med]}(x), x \in grid in A.$	Tortosa	00.5	40.8

Case studies analysis

The following four storm periods during solar cycle 24 have been analyzed: March 2012 (Case Study 1), March 2015 (Case Study 2), May 2015 (Case Study 3), and September 2017 (Case Study 4).

MUF(3000) hourly values have been used to test the method, comparing measurements and predictions at the **test stations** of **San Vito (17.8°E, 40.6°N)** and **Fairford (-01.5°E, 51.7°N)** (see figures in this panel). The results have been compared also on the basis of the RMSE values obtained between predicted and measured MUF(3000) values at the test stations (see **Table 2**).

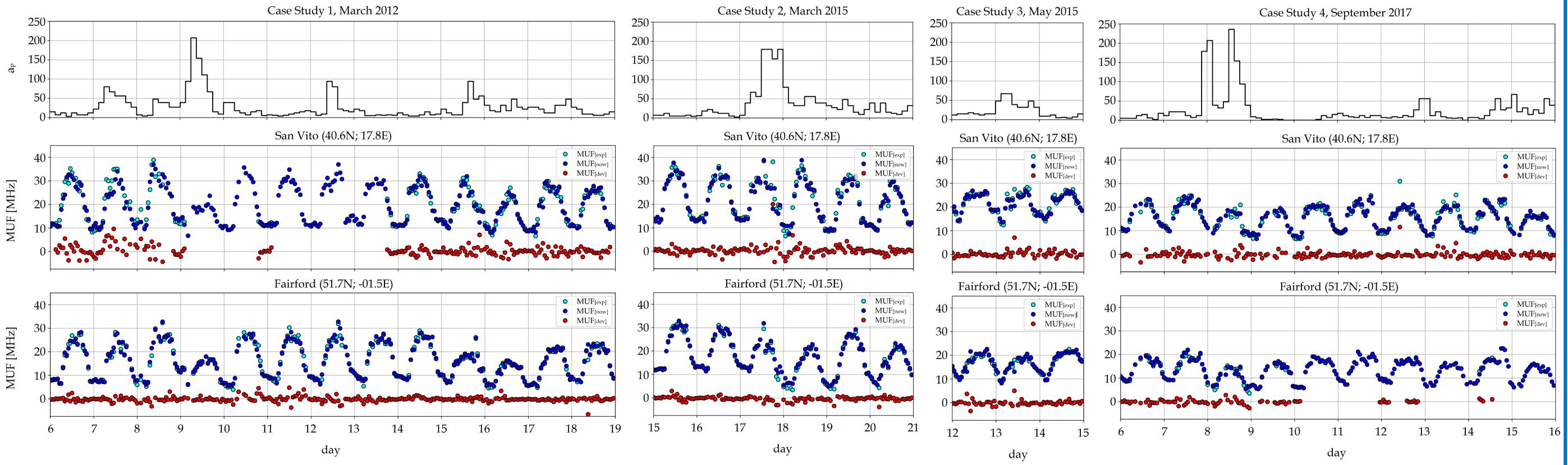
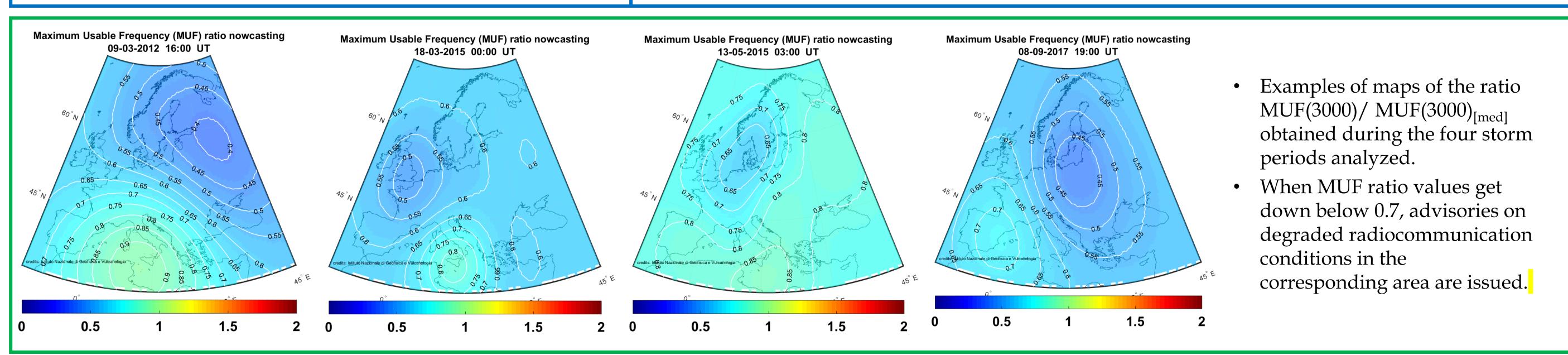


Table 2	San Vito		Fairford		
	RMSE (MHz)	n	RMSE (MHz)	n	
Case Study 1	2.15	184	1.08	291	
Case Study 2	2.27	138	0.98	127	
Case Study 3	1.36	63	1.13	62	
Case Study 4	1.44	188	0.85	86	
Total	1.90	573	1.03	566	

Results and Conclusions

- Storm periods represent the conditions of greater interest for Space Weather applications, and those
 more critical in terms of accuracy of models to reproduce ionospheric features.
- On the basis of the results reported in **Table 2**, the MUF(3000) accuracy achieved by this method during storm periods can be assessed as quite good, being the corresponding RMSE values at the test stations often less than 2 MHz, when computed over a whole storm period.
- RMSE values obtained at Fairford station result generally lower than those at San Vito, because of its
 proximity to the Chilton station, where measurements are assimilated.



Aknowledgements

The authors thank Leibniz institute of Atmospheric Physics - Field station for Juliusruh data; the PECASUS partners: Warsaw (SRC), Dourbes (STCE); the Lowell DIDBase through (http://giro.uml.edu/) and DIAS (http://dias.space.noa.gr:8080/) to provide ionospheric data; the Prediction center sites: NOAA/National GIRO Weather Service (http://www.nwra.com/spawx/list27do.html); Radio Weather (http://www.ips.gov.au/Geophysical/3/1); and Space Services Solar Influences Data Analysis Center <u>http://sidc.oma.be/products/meu/</u> for providing the data. The Rome and Gibilmanna data are provided by Istituto Nazionale di Geofisica e Vulcanologia (http://www.eswua.ingv.it/).