

Statistically Identifying Systematics from Far-side Acoustic Images

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Outline

- **Why are Far-Side Acoustic Images important?**

- **Project Overview:** (HSWO2R)

“Reliably Inferring the Sun’s Far-Side Magnetic Flux for Operations Using Time-Distance Helioseismology”

- ◆ Far-Side Acoustic Maps
- ◆ Establishing a Relationship between EUV and Magnetic Field Images
- ◆ Calibrating using a Flux Transport Model



- **Statistical Analysis So Far**

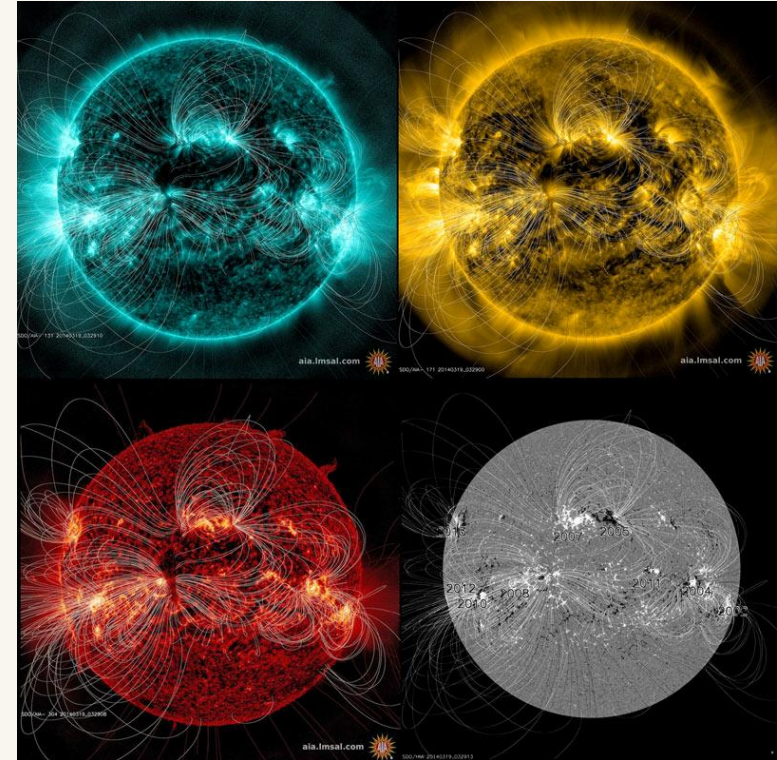
- ◆ Annual Variation
- ◆ Average Background
- ◆ Success Rate:

Acoustic Far-Side AR Detections Compared to *STEREO* 304 Å Far-Side AR Observations

Why are Far-Side Acoustic Images Important?

Current Far-Side Solutions:

- Flux Transport Models
 - ◆ Signed Flux
 - ◆ Effectively progress observed near-side flux around to far side
 - ◆ Incapable of incorporating:
 - ★ Growth of existing AR
 - ★ New flux emergence
- *STEREO/EUVI* Observations
 - ◆ Quantitative relationship btw 304 Å emission and magnetic flux → proxy (Ugarte-Urra et al. 2015)
 - ◆ No guarantee of future observations
- Current Helioseismic Imaging
 - ◆ Acoustic Holography (Lindsey & Braun 2000; Braun & Lindsey 2001)
 - ◆ Time-Distance (Zhao 2007)
 - ◆ Often give spurious signals
 - ◆ Not calibrated into magnetic flux



Courtesy of the SDO Gallery

Why are Far-Side Acoustic Images Important?

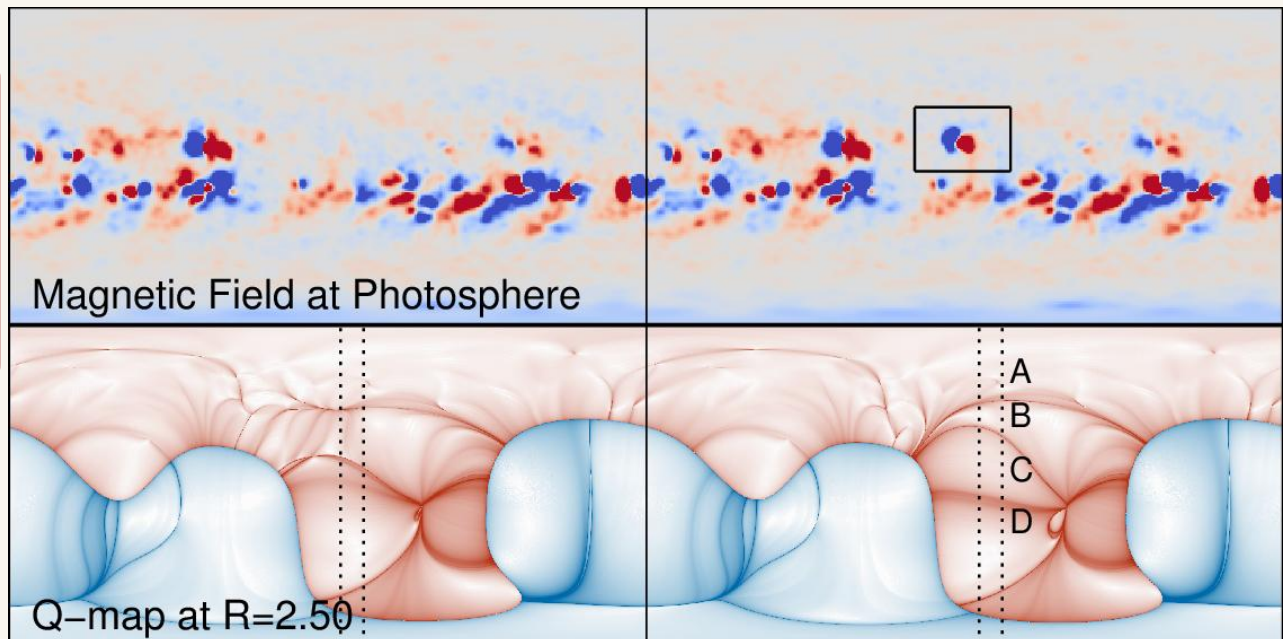
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Top: FTM data-assimilation model for 17-Feb-2014 12:00 UT (left) with an artificial far-side “emerged” AR (right)

Bottom: Q-maps calculated using above synchronic flux maps

Courtesy of M. Derosa and Y. Liu

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Current Far-Side Solutions:

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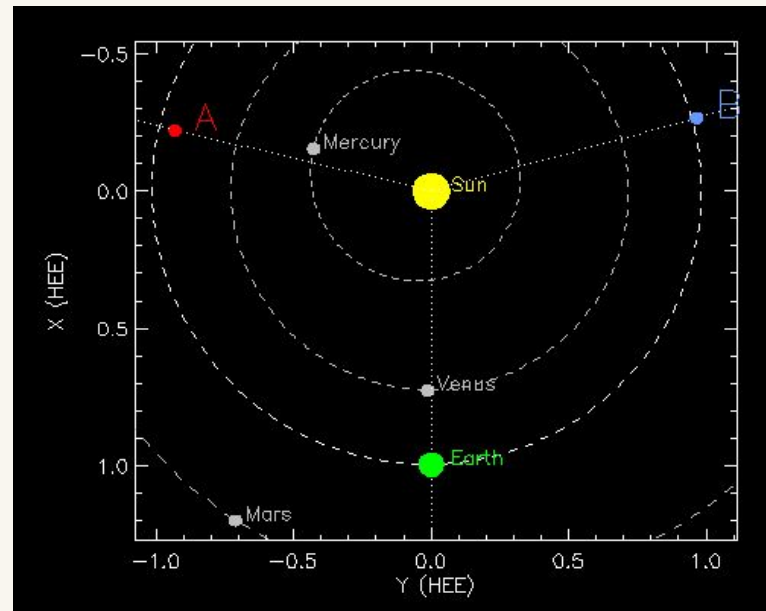
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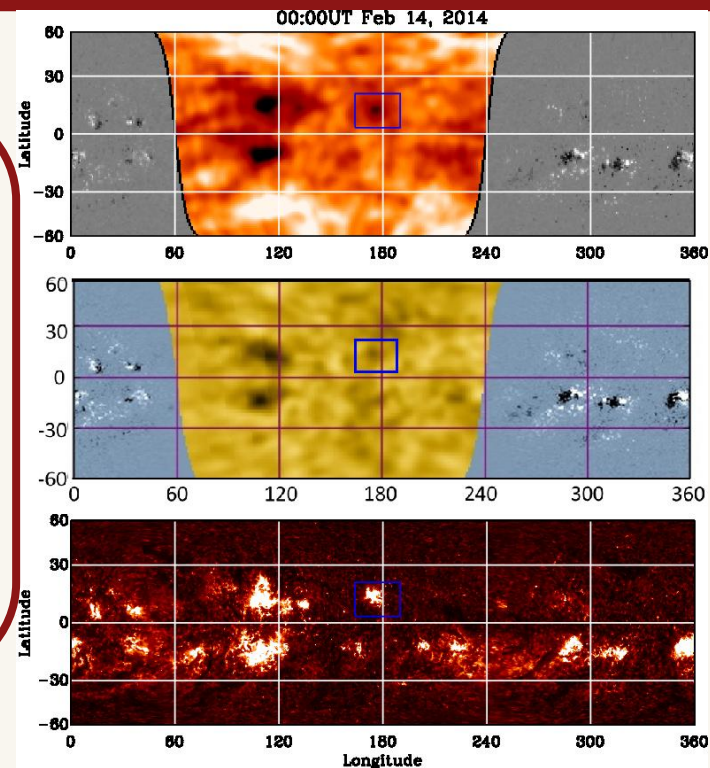
Positions of STEREO A and B for 24-Oct-2018 23:00 UT
Courtesy of the STEREO Science Center

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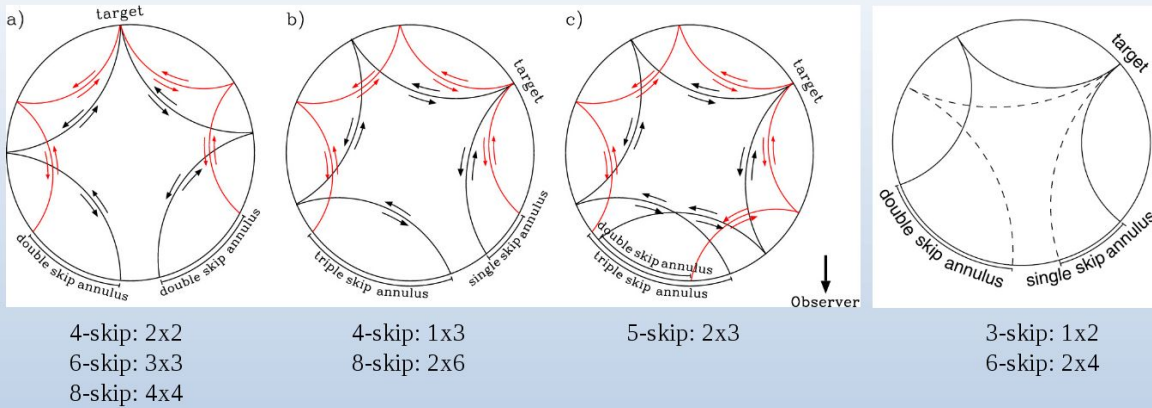
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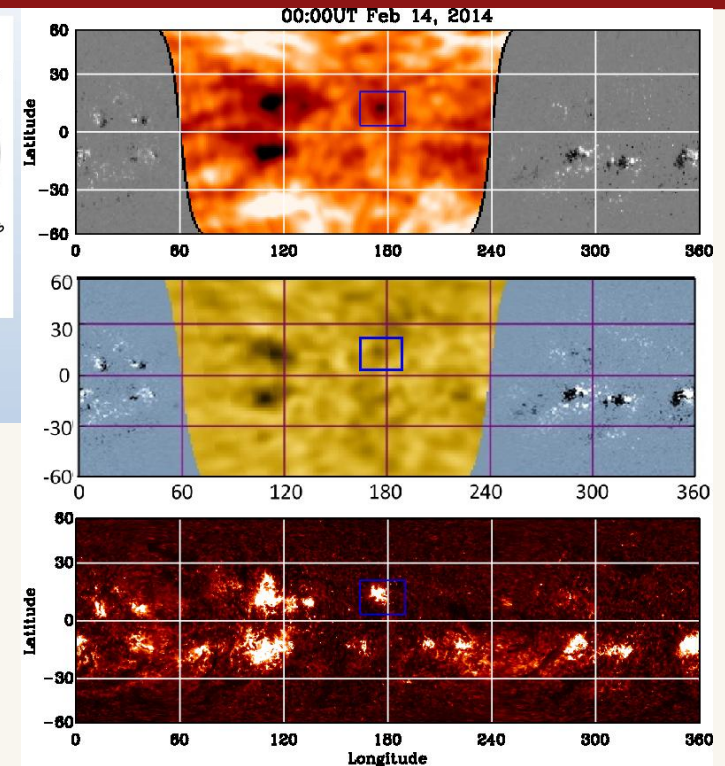
Top: Far-side travel-time acoustic map with near-side HMI magnetic field
Middle: Far-side holography acoustic map with near-side HMI magnetic field
Bottom: Synoptic map using *SDO/AIA* and *STEREO/EUVI* 304 \AA observations
Courtesy of J. Zhao

Project Overview: Far-Side Acoustic Maps



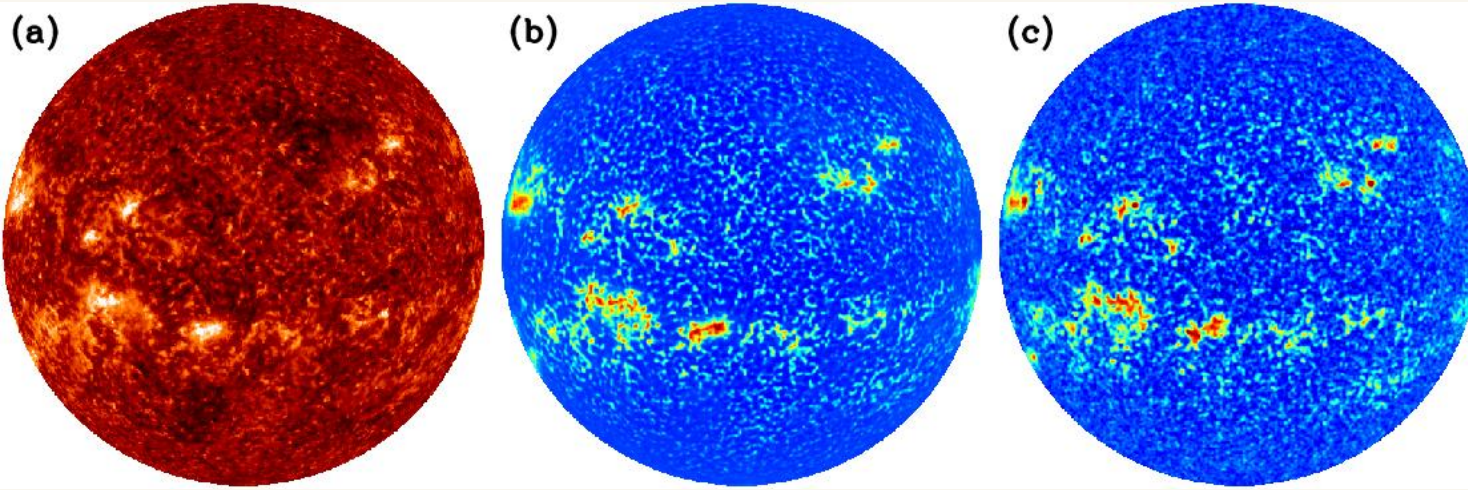
Courtesy of J. Zhao

- New Far-Side Time-Distance Technique
 - ◆ Takes advantage of helioseismic geometry for different “skip” waves
 - ◆ 14 total sets of individual calculations
 - ◆ Expanded total wave-travel distances beyond 360°



Top: Far-side travel-time acoustic map with near-side HMI magnetic field
Middle: Far-side holography acoustic map with near-side HMI magnetic field
Bottom: Synoptic map using SDO/AIA and STEREO/EUVI 304 Å observations
Courtesy of J. Zhao

Project Overview: *Deep-Learning*



For 1-Aug-2013 00:00 UT

(a) AIA 304Å image

(b-c) HMI magnetic flux data & approximated DNN result

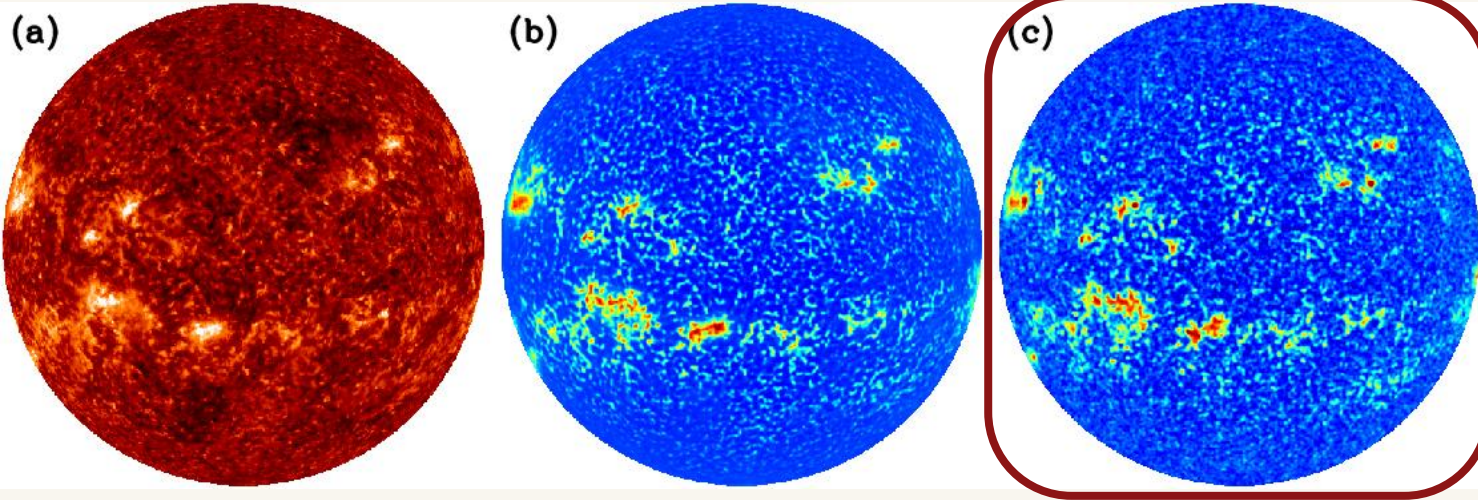
Courtesy of R. Chen

- Deep Neural Network
- 3 sets of kernels
- 8 years of training/test data
- Train on 2700 pairs of AIA 304 Å and HMI unsigned magnetic flux images

**BETWEEN (b) AND (c)
WHICH IS THE REAL HMI DATA
AND WHICH IS PRODUCED BY THE DNN?**

See e-Poster by R. Chen today (Thursday)

Project Overview: *Deep-Learning*



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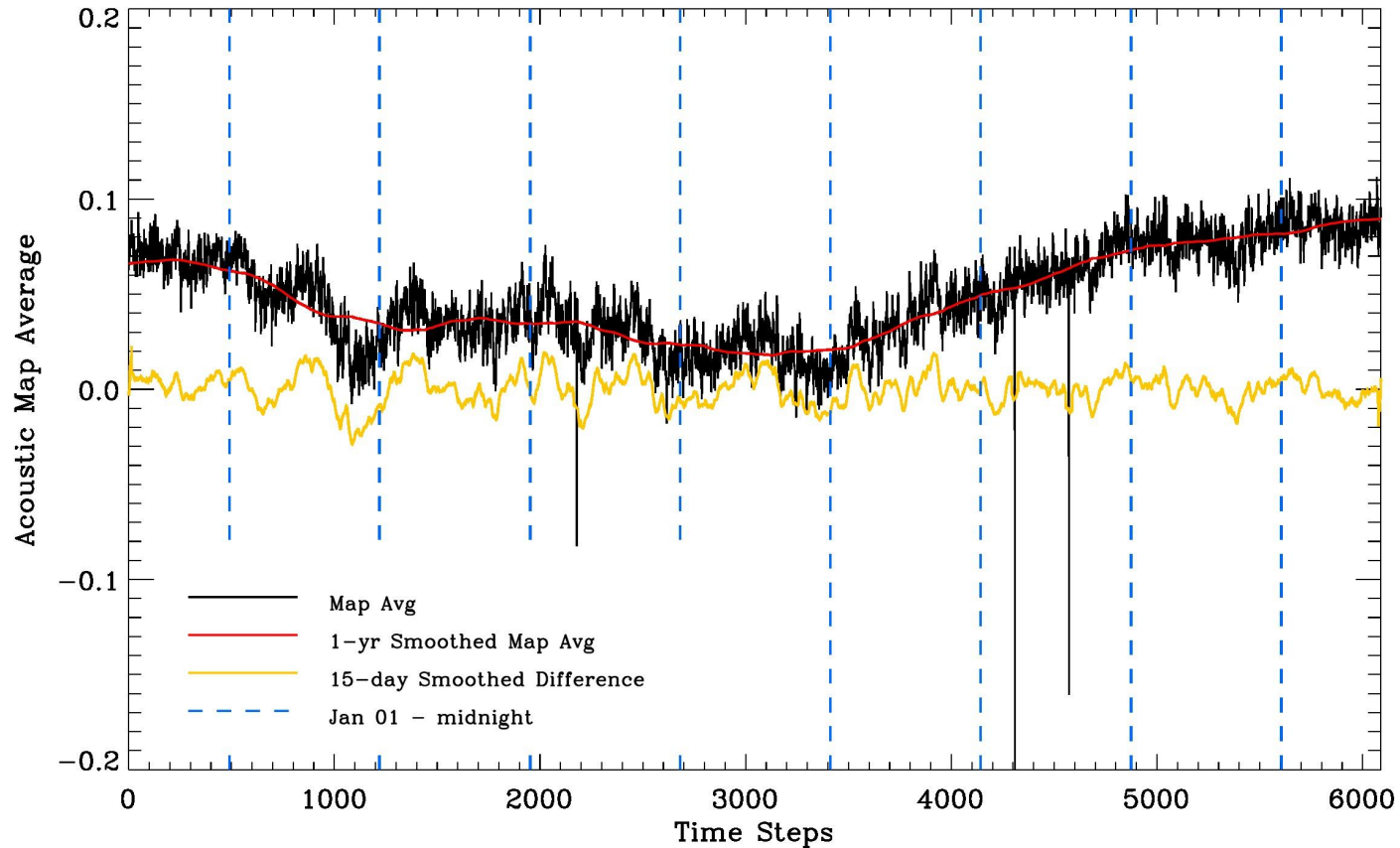
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Project Overview: *Calibrating with FTM*

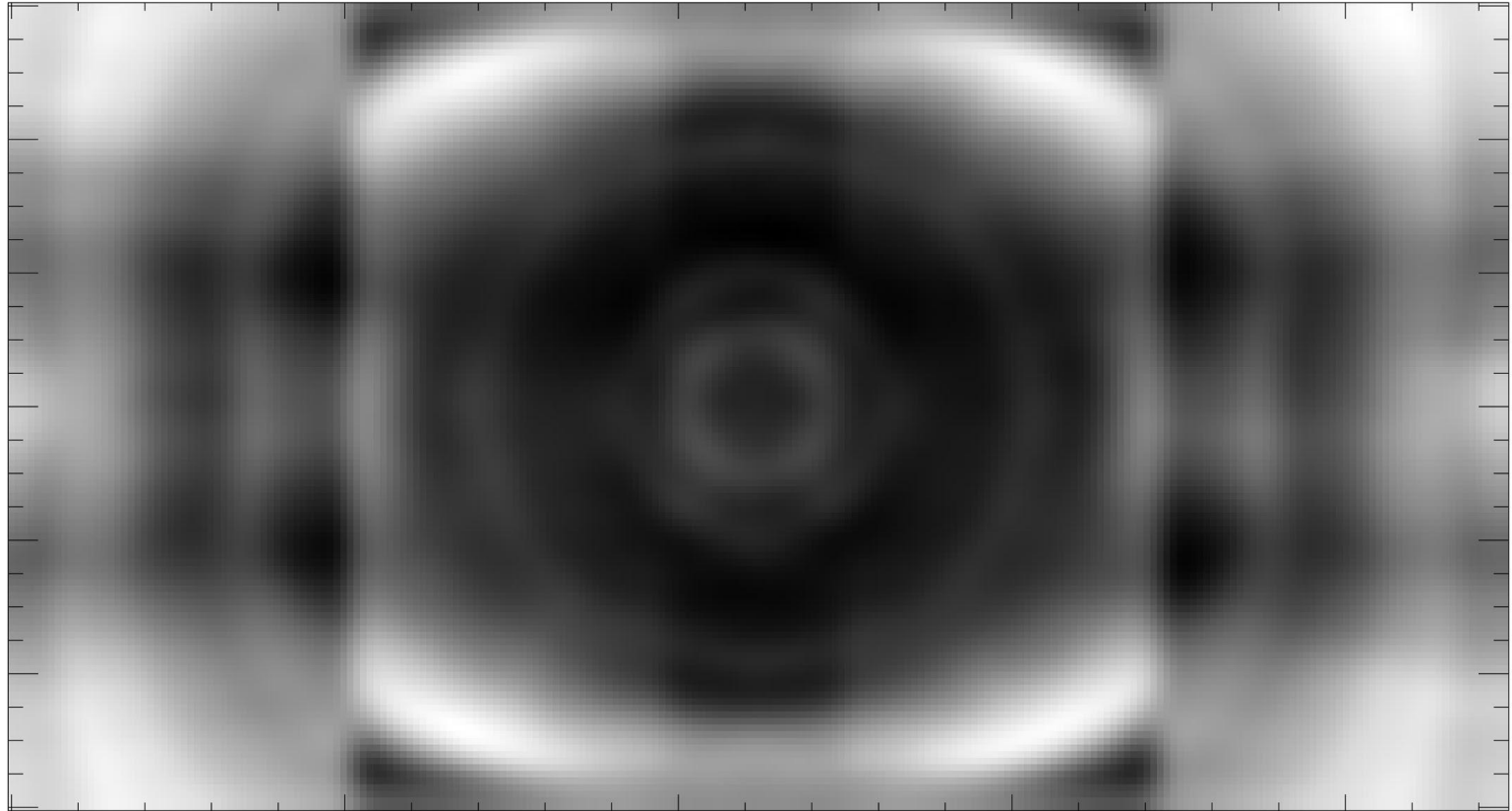
- Data-Assimilated Flux Transport Model (Schrijver & DeRosa 2003)
- Pro: Retains magnetic field polarity
- Con: Cannot model growth of existing AR or emergence of new AR
- Select at least 10 “perfectly modeled” ARs
 - ◆ Selection based on predicted total flux and position compared to region reappearance on the eastern limb
 - ◆ For selected regions, fit modeled total flux VS total acoustic travel-time shifts
 - ◆ Establish empirical relation
 - ◆ Useful for determining some uncertainty measures

Statistical Analysis: *Annual Variation*



Statistical Analysis: *Average Background*

Average Background – 2012



Statistical Analysis: *Success Rate*

Based on 3 months of statistics (Jan - Mar, 2014)

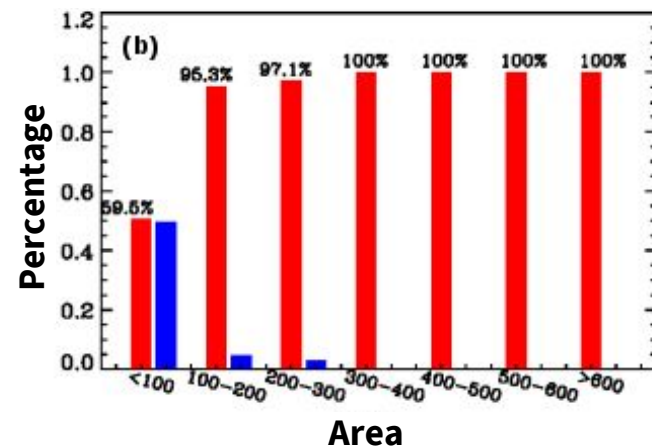
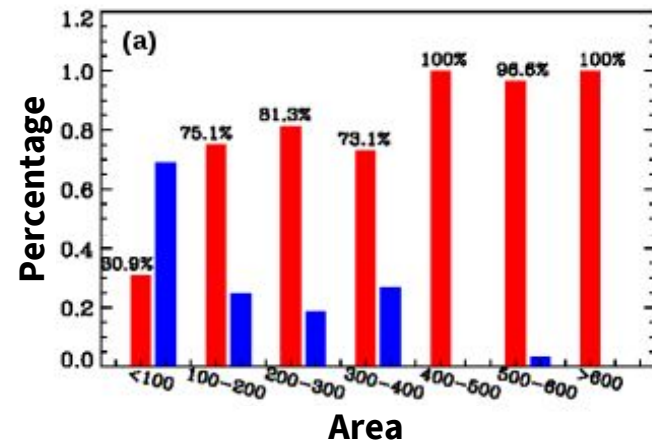
(a) For far-side ARs observed in *STEREO* 304 Å:

- ARs >400 units (1 unit = 1.4 μ hemispheres): nearly 100% detectable in acoustic images
- For ARs between 100 and 400 units: 75% are detectable in acoustic images

(b) For ARs detected in far-side acoustic images:

- ARs >100 units: nearly 100% correspond to observed *STEREO* 304 Å AR

Red: true positive (top and bottom)
Blue: false negative (top) and false positive (bottom)
Courtesy of J. Zhao



Summary

- Far-side magnetic flux maps are important for improved coronal and space weather modeling
- We are working on a new far-side magnetic flux product that uses machine learning and statistical analysis:
 - ◆ Link magnetic flux \rightarrow EUV 304 Å \rightarrow acoustic maps
 - ◆ Calibrate with a data-assimilated flux transport model
- Currently: Using statistical analysis to prepare far-side acoustic maps for use with machine learning method
- Statistical Results so far:
 - ◆ Negligible annual variation!!
 - ◆ Cannot subtract an average background without biasing data
 - ◆ For ARs ≥ 100 pixels, excellent agreement between acoustic map detections and EUV 304 Å images
- Next: Statistical study of hemispheric differences and progress on machine learning